

Exercise sheet 3b: Additional exercises

1. Proper distance, area and volumes.

Consider the metric

$$ds^2 = -(1 - Ar^2)^2 dt^2 + (1 - Ar^2)^2 dr^2 + r^2(d\vartheta^2 + \sin^2 \vartheta d\phi^2)$$

- i) Calculate the proper distance from $r = 0$ to $r = R$.
- ii) Calculate the area of a sphere with coordinate radius R .
- iii) Calculate the three-volume of a sphere with coordinate radius R .
- iv) Calculate the four-volume of a sphere with coordinate radius R bounded by two $t = \text{const}$ planes separated by time difference T .

2. Newtonian gravity as a spacetime phenomenon

Since Newtonian gravity is a special case of Einstein gravity, it should be possible to replace the Newtonian gravitational force by a deformation of Minkowski space. Confirm that the metric describing gravitational effects in the Newtonian limit can be chosen as

$$ds^2 = (1 + 2\Phi/c^2) c^2 dt^2 - (1 - 2\Phi/c^2) dl^2 \quad (1)$$

by showing that the resulting action of a point particle is equivalent to the the one using the standard Lagrangian

$$L = \frac{1}{2}mv^2 - m\Phi = T + V \quad (2)$$

in the limit $v/c \rightarrow 0$. Here, Φ is the Newtonian gravitational potential and $dl^2 = dx^2 + dy^2 + dz^2$ is the Euclidean line-element.

3. Maximal velocity of a comet

A comet starts at $r \rightarrow \infty$, approaches a star of mass M and disappears again to $r \rightarrow \infty$. What is the maximal velocity v of the comet measured by a stationary observer at the radius R of nearest approach?

(2 ways of solution: 1. find v as function of R, l and b , connect then l to b and R . 2. Use the orthonormal basis associated with the observer.)