## Exercise sheet 3

## Hartle 7-14.

Newton's law becomes in relativistic notation

$$\mathbf{F} = \frac{\mathrm{d}\mathbf{p}}{\mathrm{d}\tau} \,.$$

Calculated  $\mathbf{F} \cdot \mathbf{u}$  taking into account that the mass *m* changes. Why do we set in general m = const.?

## Hartle 7-14.

Consider the metric

$$ds^{2} = -(1 - Ar^{2})^{2}dt^{2} + (1 - Ar^{2})^{2}dr^{2} + r^{2}(d\vartheta^{2} + \sin^{2}\vartheta d\phi^{2})$$

i) Calculate the proper distance from r = 0 to r = R.

ii) Calculate the area of a sphere with coordinate radius R.

iii) Calculate the three-volume of a sphere with coordinate radius R.

iv) Calculate the four-volume of a sphere with coordinate radius R bounded by two t=const planes separated by time difference T.

## Spherical coordinates II.

Calculate for spherical coordinates  $x = (r, \vartheta, \phi)$  in  $\mathbb{R}^3$  the gradient, divergence, and the Laplace operator. Note that one uses usually normalized unit vectors in case of a diagonal metric: this corresponds to a rescaling of vector components  $V^i \to V^i / \sqrt{g_{ii}}$  (no summation in *i*) or basis vectors.