

# NTNU Trondheim, Institutt for fysikk

## Examination for FY3452 Gravitation and Cosmology

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Possible languages for your answers: *Bókmal, Castellano, English, Nynorsk.*

Allowed tools: *Pocket calculator, mathematical tables*

Some formulas can be found at the end of p.2.

### 1. Hyperbolic plane $H^2$ .

The line-element of the Hyperbolic plane  $H^2$  is given by

$$ds^2 = y^{-2}(dx^2 + dy^2) \quad \text{and} \quad y \geq 0.$$

- Write out the geodesic equations and deduce the Christoffel symbols  $\Gamma^a_{bc}$ . (4 pts)
- Calculate the Riemann (or curvature) tensor  $R^a_{bcd}$  and the scalar curvature  $R$ . (4 pts)

### 2. Kerr metric.

The metric outside a spherically symmetric mass distribution with mass  $M$  and angular momentum  $J$  is given by

$$ds^2 = - \left[ 1 - \frac{2Mr}{\rho^2} \right] dt^2 - \frac{4Mar \sin^2 \vartheta}{\rho^2} d\phi dt + \frac{\rho^2}{\Delta} dr^2 + \rho^2 d\vartheta^2 + \left[ r^2 + a^2 + \frac{2Mra^2 \sin^2 \vartheta}{\rho^2} \right] \sin^2 \vartheta d\phi^2,$$

with

$$a = \frac{J}{M}, \quad \rho^2 = r^2 + a^2 \cos^2 \vartheta, \quad \Delta = r^2 - 2Mr + a^2.$$

- Find the outer boundary of the ergosphere, i.e. the surface enclosing the region where no stationary observers are possible in the Kerr metric. (3 pts)
- Find the two horizons of the Kerr metric. (1.5 pt)
- Determine the smallest possible unstable circular orbit of a massive particle for  $J = 0$ . (Hint: Consider the effective potential  $V_{\text{eff}}$ .) (6 pt)

### 3. Scalar fields in FLRW metric.

Consider a scalar field  $\phi$  with potential  $V$

$$\mathcal{L} = \frac{1}{2} g^{ab} \nabla_a \phi \nabla_b \phi + V(\phi)$$

in a flat FLRW metric,

$$ds^2 = -dt^2 + a^2(t) \left[ \frac{dr^2}{1 - kr^2} + r^2(\sin^2 \vartheta d\phi^2 + d\vartheta^2) \right].$$

- Derive the equation of motions for  $\phi$ . (4 pts)
- Derive the energy-momentum tensor for  $\phi$ . (3 pts)

- c. Derive the equation of state  $w = P/\rho$  for  $\phi$  assuming that the field  $\phi$  is uniform in space,  $\phi(t, \vec{x}) = \phi(t)$ . (2 pts)
- d. Scalar fields are often used as models for inflation. Give *one* necessary condition that  $\phi$  can drive inflation. (1 pt)

#### 4. Killing vectors.

Consider Minkowski space of special relativity,

$$ds^2 = -dt^2 + dx^2 + dy^2 + dz^2.$$

- a. Find all ten Killing vectors and name the conserved symmetries and conserved quantities. (4.5 pt)

#### 5. Radiation from a particle in a gravitational field.

An electron is released at the position  $r \gg 2M$  in the gravitational field of a point mass  $M$  and moves thereafter on a geodesics. Give either a *short, simple* argument why

- a. the electron does not emit radiation. (1 pt)
- and
- b. the electron does emit radiation. (1 pt)
- or
- c. decide which one of the alternatives is correct and explain why. (2 pt)

**Some formula:** Signature of the metric  $(-, +, +, +)$ .

$$T^{ab} = (\rho + P)u^a u^b + P g^{ab}$$

$$\nabla_i \xi_j + \nabla_j \xi_i = 0$$

$$\ddot{x}^c + \Gamma_{ab}^c \dot{x}^a \dot{x}^b = 0$$

$$\frac{2}{\sqrt{|g|}} \frac{\delta S_m}{\delta g^{ab}} = T_{ab}$$

$$R^a_{bcd} = \partial_c \Gamma^a_{bd} - \partial_d \Gamma^a_{bc} + \Gamma^a_{ec} \Gamma^e_{bd} - \Gamma^a_{ed} \Gamma^e_{bc},$$

$$\delta \sqrt{|g|} = \frac{1}{2} \sqrt{|g|} g^{ab} \delta g_{ab}$$