#### Exercise sheet 2

### 1. Scalar product of time-like vectors.

Show that the scalar product of two time-like vectors can be expressed as  $\mathbf{a} \cdot \mathbf{b} = ab \cosh \eta$ , where  $\eta$  is the rapidity connecting the two frames where  $a^{\mu} \equiv (a, \mathbf{0})$  and  $\tilde{b}^{\mu} \equiv (b, \mathbf{0})$  are valid.

Performing a Lorentz transformation from the frame A where  $\mathbf{a} = (a, \mathbf{0})$  to the frame B gives

$$\tilde{a} = (a \cosh \eta, a \sinh \eta, 0, 0)$$

where we assumed that the boost is in  $x^1$  direction. Taking then the scalar product results in

$$\mathbf{a} \cdot \mathbf{b} = \eta_{\mu\nu} \tilde{a}^{\mu} \tilde{b}^{\nu} = ab \cosh \eta.$$

Remarks: i) The result is in line with the usual  $\mathbf{a} \cdot \mathbf{b} = ab \cos \alpha$  and our interpretation of  $\eta$  as imaginary rotation angle. ii) Since  $\cosh \eta \geq 1$ , two time-like vectors cannot be orthogonal – as expected for two elements of an one-dimensional subspace.

# 2. Action of free relativistic particle.

Consider  $S = \alpha \int d\tau$  as action for a free relativistic particle.

- a.) Determine the constant  $\alpha$  requiring the correct non-relativistic limit.
- b.) Does a classically allowed path maximise or minimise the action?
- a.) We ask that the action has the correct non-relativistic limit/ Then

$$S_0 = \alpha \int_a^b ds = \alpha \int_a^b dt \sqrt{1 - v^2} = \int_a^b dt \left( -m + \frac{1}{2} m v^2 + \mathcal{O}(v^4) \right) , \tag{1}$$

if we set  $\alpha = -m$ . The mass m corresponds to a potential energy in the non-relativistic limit and has therefore a negative sign in the Lagrangian. Moreover, a constant drops out of the equations of motion, and thus the term -m can be omitted in the non-relativistic limit.

b.) Compare with Fig. 3.1 in the script and the corresponding discussion.

### 3. Uniformly accelerated observer.

Consider a particle moving on the x axis along a world-line parametrised by

$$t(\sigma) = \frac{1}{a} \sinh \sigma, \qquad x(\sigma) = \frac{1}{a} \cosh \sigma.$$

- a.) Find the connection between  $\sigma$  and proper-time  $\tau$ ; express the world-line as function of  $\tau$ .
- b.) Calculate the four-velocity  $u^{\alpha}$  and the three-velocity  $v^{1}$  of the particle. Check their normalisation.

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- c.) Calculate the four-acceleration  $a^{\alpha}$  of the particle.
- d.) Draw a spacetime diagram including  $x^{\mu}(\sigma)$ .
- a.) We form the differential,

$$\mathrm{d}\tau^2 = \mathrm{d}s^2 = \mathrm{d}t^2 - \mathrm{d}x^2 = \left(\frac{1}{a}\cosh\sigma\mathrm{d}\sigma\right)^2 - \left(\frac{1}{a}\sinh\sigma\mathrm{d}\sigma\right)^2 = \left(\cosh^2\sigma - \sinh^2\sigma\right)\left(\frac{\mathrm{d}\sigma}{a}\right)^2 = \left(\frac{\mathrm{d}\sigma}{a}\right)^2.$$

Setting  $\tau = \sigma/a$ , i.e. choosing the constant to zero, it follows

$$t(\tau) = \frac{1}{a}\sinh(a\tau), \qquad x(\tau) = \frac{1}{a}\cosh(a\tau).$$

b.) We differentiate w.r.t.  $\tau$ ,

$$u^{0} = \frac{\mathrm{d}t}{\mathrm{d}\tau} = \cosh(a\tau), \qquad u^{1} = \frac{\mathrm{d}x}{\mathrm{d}\tau} = \sinh(a\tau),$$

Checking the normalisation gives  $\boldsymbol{u} \cdot \boldsymbol{u} = \cosh^2(a\tau) - \sinh^2(a\tau) = 1$ . The three-velocity is

$$v^1 = \frac{\mathrm{d}x}{\mathrm{d}t} = \frac{\mathrm{d}x}{\mathrm{d}\tau} \frac{\mathrm{d}t}{\mathrm{d}\tau} = \frac{u^1}{u^0} = \tan(a\tau),$$

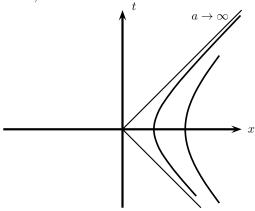
and thus bounded by one,  $|v| \le 1$ : The particle has the velocity v = -c for  $\to -\infty$ , decelerates to v(t=0) = 0, followed by acceleration to v = c for  $\to \infty$ .

c.) We differentiate another time w.r.t.  $\tau$ ,

$$a^{0} = \frac{\mathrm{d}u^{0}}{\mathrm{d}\tau} = a \sinh(a\tau), \qquad a^{1} = \frac{\mathrm{d}u^{1}}{\mathrm{d}\tau} = a \cosh(a\tau),$$

The norm is  $|\boldsymbol{a} \cdot \boldsymbol{a}|^{1/2} = -a$ , thus the acceleration is constant.

d.) The worldline of an uniformly accelerated observers O; in the limit  $a \to \infty$ , the world-line become a triangle. Note the presence of 2 horizonts: One half of Minkowski space cannot influence the O, while O cannot influence half of Minkowski space.



Extra: If you want some additional training, you can solve the inverse problem: Determine the trajectory of a particle which is uniformly accelerated in its rest-frame.

In the rest-frame of an uniformly accelerated observer, the four-acceleration is given by  $a^{\alpha} = \ddot{x}^{\alpha} = (0, \mathbf{a})$  with  $|\mathbf{a}| = a = \text{const.}$  We can convert this condition into a covariant form, writing

$$\eta_{\alpha\beta}\ddot{x}^{\alpha}\ddot{x}^{\beta} = -a^2. \tag{2}$$

In order to determine the trajectory  $x^{\alpha}(\tau)$  of the accelerated observer, it is convenient to change to light-cone coordinates,

$$u = t - x$$
 and  $v = t + x$ .

(We will suppress the transverse coordinates y and z.) Forming the differentials du and dt, we see that the line element in the new coordinates is  $ds^2 = dudv$ . The normalisation condition  $\eta_{\alpha\beta}\dot{x}^{\alpha}\dot{x}^{\beta} = 1$  of the four-velocity becomes therefore  $\dot{u}\dot{v} = 1$ , while the acceleration equation (2) results in  $\ddot{u}\ddot{v} = -a^2$ . Differentiating then  $\dot{u} = 1/\dot{v}$ , we obtain  $\ddot{u} = -\ddot{v}/\dot{v}^2$  or

$$\frac{\ddot{v}}{\dot{v}} = \pm a.$$

Integrating results in

$$v(\tau) = \frac{A}{a} \exp(a\tau) + C$$

and, using  $\dot{u} = 1/\dot{v}$ , in

$$u(\tau) = -\frac{1}{4a} \exp(-a\tau) + D.$$

Going back to the original Cartesian coordinates, we obtain

$$t(\tau) = \frac{1}{a}\sinh(a\tau)$$
 and  $x(\tau) = \frac{1}{a}\cosh(a\tau)$ ,

where we set the integration constants A = 1 and C = D = 0, which selects the trajectory with t(0) = 0 and x(0) = 1/a.

# 4. Infinitesimal Lorentz transformation.

Symmetry transformations form groups; continuous transformations in physics depend analytically on their parameters (e.g. as  $\cos \vartheta$  and  $\sin \vartheta$  on the rotation angle  $\vartheta$ ). An element g of such a group (called "Lie group") can be therefore expanded as a power series,

$$g(\vartheta) = 1 + \sum_{a=1}^{n} i\vartheta^{a} T^{a} + \mathcal{O}(\vartheta^{2}) \equiv 1 + i\vartheta^{a} T^{a} + \mathcal{O}(\vartheta^{2}). \tag{3}$$

The linear transformation in the arbitrary direction  $\vartheta^a$  is called an infinitesimal transformation, the  $T^a$  the (infinitesimal) generators of the transformation. The generators  $T^a$  can be obtained by differentiation,  $T^a = -\mathrm{i} \,\mathrm{d} g(\vartheta)/\mathrm{d} \vartheta^a|_{\vartheta=0}$ . Conversely, analyticity implies that the group element  $g(\vartheta)$  (more precisely, thus connected to the unity element) can be obtained by exponentiation,

$$g(\vartheta) = \lim_{n \to \infty} \left[ 1 + i\vartheta^a T^a / n \right]^n \tag{4}$$

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- a.) Calculate the generators of Lorentz transformations.
- b.) Determine their "Lie algebra", i.e. calculate the real numbers  $f^{abc}$  called structure constants in

$$[T^a, T^b] = if^{abc}T^c. (5)$$

Applied to the finite boost  $B_x(\eta)$  along the x direction given in (??) we find as generator  $K_x$ 

and similarly for the other two boosts. The 4-dim. generators of rotations are obtained simply by adding (1,0,0,0) as zeroth colum and raw to the known 3-dim. rotations, e.g.

$$R_{z}(\alpha) = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & \cos \alpha & \sin \alpha & 0 \\ 0 & -\sin \alpha & \cos \alpha & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}, \qquad J_{z}(\alpha) = -i \frac{\partial R_{z}(\alpha)}{\partial \alpha} \Big|_{\alpha=0} = -i \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix}.$$

$$(7)$$

Calculating then their commutation relations, one finds

$$[J_i, J_i] = i\varepsilon_{ijk}J_k, \tag{8a}$$

$$[J_i, K_i] = i\varepsilon_{ijk}K_k, \tag{8b}$$

$$[K_i, K_j] = -i\varepsilon_{ijk}J_k. \tag{8c}$$

You should recognise in the first relation the anticommutation relation of the angular momentum operators; thus angular momentum are the generator of rotations. Moreover, their anticommutation relation are closed (i.e. contain no  $K_i$ ). Thus rotations from a closed subgroup of the Lorentz group. This is not true for the boosts.

## Remarks:

- The structure constant  $\varepsilon_{ijk}$  are those of SU(2). Introducing the linear combinations  $J^{\pm} = (J \pm iK)/2$ , one finds that their mixed commutators are zero, while both  $J^{\pm}$  form a SU(2) group. Thus SO(1,3)~ SU(2) $\otimes$ SU(2).
- You may have encountered SU(2) when you discussed rotations of Pauli spinors. Thus in addition of Lorentz transformations acting on tensors, we could construct Lorentz transformations acting on spinors using  $SU(2) \otimes SU(2)$ .
- Writing  $U = \exp(iT^a\vartheta^a)$ , we factored out an i. Then generators like J are Hermetian (i.e.  $(J_z)^\dagger = (J_z)^{\mathrm{T},*} = J_z$ ) and U is an unitary transformation. An exception are boosts K: They cannot be implemented by unitary (finite-dimensional) matrices; physically, this is clear since the number density  $j^0$  of particles is not Lorentz invariant.