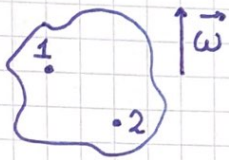


5. Rigid body equations of motion

(66)

5.1 Angular momentum and kinetic energy

Rotating rigid body:  $\vec{\omega}_1 = \vec{\omega}_2 = \vec{\omega}$

Assume rotation around a fixed point = origo = CM.

Total angular momentum (around the fixed origin):

$$\vec{L} = m_i (\vec{r}_i \times \vec{v}_i) \quad [\text{Sum convention}]$$

Rotation only $\Rightarrow \vec{v}_i = \vec{\omega} \times \vec{r}_i$ Check this, using ϵ_{ijk}

$$\Rightarrow \vec{L} = m_i \vec{r}_i \times (\vec{\omega} \times \vec{r}_i) = m_i [\vec{\omega} r_i^2 - \vec{r}_i (\vec{r}_i \cdot \vec{\omega})]$$

Components of $\vec{L}$:

$$\begin{aligned} L_x &= m_i [\omega_x r_i^2 - x_i (x_i \omega_x + y_i \omega_y + z_i \omega_z)] \\ &= m_i [\omega_x (r_i^2 - x_i^2) - \omega_y x_i y_i - \omega_z x_i z_i] \end{aligned}$$

On the general form (with j and k denoting components)

$$L_j = I_{jk} \omega_k ; \quad I_{jk} = (\text{moment of}) \text{ inertia tensor}$$

with elements $I_{xx} = m_i (r_i^2 - x_i^2)$, $I_{xy} = -m_i x_i y_i$ etc.

Continuous mass distribution: $m_i \rightarrow dm = \rho dV$; $\sum_i \rightarrow \int_V$

$$\Rightarrow I_{xx} = \int_V \rho(\vec{r}) (r^2 - x^2) dV, \quad I_{xy} = - \int_V \rho(\vec{r}) xy dV$$

$$\begin{aligned} \text{Compact: } I_{jk} &= \int_V \rho(\vec{r}) (r^2 \delta_{jk} - x_j x_k) dV \quad [xyz \rightarrow x_1 x_2 x_3] \\ \vec{L} &= \vec{I} \vec{\omega} \end{aligned}$$

Clearly $\overset{\leftrightarrow}{I}$ is symmetric, $I_{jk} = I_{kj}$, with only real elements, i.e., $\overset{\leftrightarrow}{I}$ is hermitean (or self-adjoint, cf. QM). Then $\overset{\leftrightarrow}{I}$ can always be diagonalized, i.e., the body set of coords. can be transformed into a coord. system where $\overset{\leftrightarrow}{I}$ is diagonal:

$$\overset{\leftrightarrow}{I} = \begin{pmatrix} I_1 & 0 & 0 \\ 0 & I_2 & 0 \\ 0 & 0 & I_3 \end{pmatrix} = \text{principal moment of inertia tensor}$$

x_1, x_2, x_3 = principal axes

$L_j = I_j \omega_j$ = angular momentum components ($j=1,2,3$)
(No sum...!)

Ex: Symmetrical top (Snurrebass): $I_1 = I_2 \neq I_3$

Spherical top: $I_1 = I_2 = I_3 = I \Rightarrow \vec{L} = I \vec{\omega}$

Kinetic energy for rotation around fixed point:

$$\begin{aligned} T &= \frac{1}{2} m_i v_i^2 = \frac{1}{2} m_i \vec{v}_i \cdot \vec{v}_i = \frac{1}{2} m_i \vec{v}_i \cdot (\vec{\omega} \times \vec{r}_i) \\ &= \frac{1}{2} m_i \vec{\omega} \cdot (\vec{r}_i \times \vec{v}_i) \quad [\text{Check yourself}] \\ &= \frac{1}{2} \vec{\omega} \cdot \vec{L} = \frac{1}{2} \vec{\omega} \overset{\leftrightarrow}{I} \vec{\omega} = \frac{1}{2} \omega_j I_{jk} \omega_k \end{aligned}$$

If rotation around axis $\hat{n}$, i.e., $\vec{\omega} = \omega \hat{n}$:

$$T = \frac{1}{2} \omega n_j I_{jk} \omega n_k = \frac{1}{2} (n_j I_{jk} n_k) \omega^2 = \frac{1}{2} I \omega^2$$

with $I \equiv n_j I_{jk} n_k = \hat{n} \overset{\leftrightarrow}{I} \hat{n}$ = mom. of inertia around rot. axis $\hat{n}$

In coord. system with principal axes x_1, x_2, x_3 :

$$T = \frac{1}{2} I_j \omega_j^2$$

5.5 Euler equations of motion

(68)

From p.4: $\vec{\tau} = \dot{\vec{L}}_S$ (measured in the inertial frame)

Measured in the rotating system (p.64; drop index b):

$$\vec{\tau} = \dot{\vec{L}} + \vec{\omega} \times \vec{L} \quad (\text{Euler eqns. with one point fixed})$$

Assume body axes along the principal axes such that $\vec{I}$ is diagonal and $L_j = I_j \omega_j$. Then:

$$\tau_i = \dot{L}_i + \epsilon_{ijk} \omega_j L_k \quad ; \quad i=1,2,3$$

Explicit:

$$\tau_1 = I_1 \dot{\omega}_1 - \omega_2 \omega_3 (I_2 - I_3)$$

$$\tau_2 = I_2 \dot{\omega}_2 - \omega_3 \omega_1 (I_3 - I_1)$$

$$\tau_3 = I_3 \dot{\omega}_3 - \omega_1 \omega_2 (I_1 - I_2)$$

5.6 Free rotation. Precession

Assume $\vec{\tau} = 0$ and symmetrical body with $I_1 = I_2 \neq I_3$.

$$\Rightarrow I_1 \dot{\omega}_1 = \omega_2 \omega_3 (I_1 - I_3)$$

$$I_2 \dot{\omega}_2 = \omega_1 \omega_3 (I_3 - I_1)$$

$$I_3 \dot{\omega}_3 = 0$$

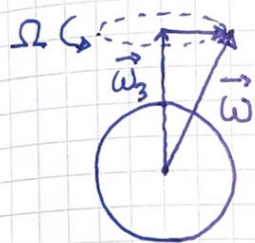
$$\Rightarrow \omega_3 = \text{const.} \quad (\text{determined by initial conditions})$$

$$\Rightarrow \dot{\omega}_1 = -\Omega \omega_2, \quad \dot{\omega}_2 = \Omega \omega_1; \quad \Omega = \frac{I_3 - I_1}{I_1} \omega_3$$

$$\Rightarrow \ddot{\omega}_1 + \Omega^2 \omega_1 = 0 \quad (\text{after eliminating } \omega_2)$$

$$\Rightarrow \omega_1(t) = A \cos \Omega t, \quad \omega_2(t) = A \sin \Omega t$$

⇒ $\vec{\omega}$ precesses around the body x_3 axis, (69)
with ~~the~~ angular frequency Ω and constant $\omega = |\vec{\omega}|$:



With $\vec{r} = 0$ both T and L^2 are conserved:

$$\left. \begin{aligned} T &= \frac{1}{2} I_1 \omega_1^2 + \frac{1}{2} I_3 \omega_3^2 \\ L^2 &= I_1^2 \omega_1^2 + I_3^2 \omega_3^2 \end{aligned} \right\} \Rightarrow A(T, L) \text{ and } \omega_3(T, L)$$

Rotating (spinning) planets: Chandler wobble)

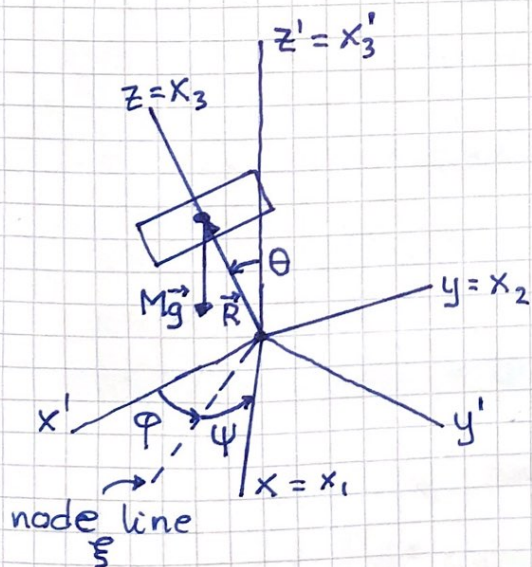
$$I_3 - I_1 \ll I_1 \Rightarrow \Omega \ll \omega_3$$

Earth: $2\pi/\omega_3 = 1 \text{ day}$; $2\pi/\Omega = 433 \text{ days}$

Mars: $2\pi/\omega_3 \approx 1 \text{ day } 40 \text{ min}$; $2\pi/\Omega = 207 \text{ days}$

[A.S. Konopliv et al, Geophysical Research Letters, Oct. 2020]

5.7 Symmetrical top with one point fixed



The top spins around the x_3 axis (a principal axis) with the origin a fixed point.

External torque around origin:

$$\vec{\tau} = \vec{R} \times M\vec{g}$$

x_j fixed in the top

x'_j fixed in space

NB: Opposite in chapter 4!!

$$\Rightarrow \vec{g} = -g \hat{x}_3^p; \quad \vec{R} = l \hat{z} = l \hat{x}_3 \quad (l = |\vec{R}|)$$

Symmetry $\Rightarrow I_1 = I_2 \neq I_3$

Kinetic energy: $T = \frac{1}{2} I_j \omega_j^2 = \frac{1}{2} I_1 (\omega_1^2 + \omega_2^2) + \frac{1}{2} I_3 \omega_3^2$

1 point fixed $\Rightarrow$ 3 independent degrees of freedom (70)

Will use the Euler angles φ, θ, ψ to construct the Lagrangian $L = T - V \Rightarrow$ Need $\omega_j(q_i, \dot{q}_i)$, i.e., expressions for the components of $\vec{\omega}$ along the body axes x_j ($j=1,2,3$). We transform from space to body by considering $\vec{\omega}'$ as 3 successive rotations with angular velocities $\omega_{\varphi}' = \dot{\varphi}$, $\omega_{\theta}' = \dot{\theta}$ and $\omega_{\psi}' = \dot{\psi}$.

$$\vec{\omega}_{\varphi}' = \dot{\varphi} \hat{x}_3' \rightarrow \begin{pmatrix} 0 \\ 0 \\ \dot{\varphi} \end{pmatrix} \Rightarrow \vec{\omega}_{\varphi} = A \vec{\omega}_{\varphi}' = B C D \vec{\omega}_{\varphi}'$$

With A from p. 61:

$$(\omega_{\varphi})_1 = \dot{\varphi} \sin \theta \sin \psi, (\omega_{\varphi})_2 = \dot{\varphi} \sin \theta \cos \psi, (\omega_{\varphi})_3 = \dot{\varphi} \cos \theta$$

$$\vec{\omega}_{\theta}' = \dot{\theta} \hat{z} \rightarrow \begin{pmatrix} \dot{\theta} \\ 0 \\ 0 \end{pmatrix} \Rightarrow \vec{\omega}_{\theta} = B \vec{\omega}_{\theta}'$$

With B from p. 60:

$$(\omega_{\theta})_1 = \dot{\theta} \cos \psi, (\omega_{\theta})_2 = -\dot{\theta} \sin \psi, (\omega_{\theta})_3 = 0$$

$$\vec{\omega}_{\psi}' = \dot{\psi} \hat{x}_3 \rightarrow \begin{pmatrix} 0 \\ 0 \\ \dot{\psi} \end{pmatrix} \Rightarrow \vec{\omega}_{\psi} = \vec{\omega}_{\psi}'$$

$$(\omega_{\psi})_1 = 0, (\omega_{\psi})_2 = 0, (\omega_{\psi})_3 = \dot{\psi}$$

We add contributions to each component ω_j :

$$\omega_1 = (\omega_{\varphi})_1 + (\omega_{\theta})_1 + (\omega_{\psi})_1 = \dot{\varphi} \sin \theta \sin \psi + \dot{\theta} \cos \psi$$

$$\omega_2 = \dot{\varphi} \sin \theta \cos \psi - \dot{\theta} \sin \psi$$

$$\omega_3 = \dot{\varphi} \cos \theta + \dot{\psi}$$

$$\Rightarrow \omega_1^2 + \omega_2^2 = \dot{\varphi}^2 \sin^2 \theta + \dot{\theta}^2$$

$$\Rightarrow T = \frac{1}{2} I_1 (\dot{\varphi}^2 \sin^2 \theta + \dot{\theta}^2) + \frac{1}{2} I_3 (\dot{\varphi} \cos \theta + \dot{\psi})^2$$

$$\text{Pot. energy: } V = M g l \cos \theta \quad [V=0 \text{ in } x'y' \text{ plane}]$$

Lagrangian: $L = T - V = L(\theta, \dot{\phi}, \dot{\theta}, \dot{\psi})$

(71)

$\Rightarrow \phi, \psi$ cyclic coords. $\Rightarrow P_\phi, P_\psi$ conserved:

$$\left. \begin{aligned} P_\psi &= \frac{\partial L}{\partial \dot{\psi}} = I_3 (\dot{\psi} + \dot{\phi} \cos \theta) = I_3 \omega_3 \equiv I_1 a \\ P_\phi &= \frac{\partial L}{\partial \dot{\phi}} = (I_1 \sin^2 \theta + I_3 \cos^2 \theta) \dot{\phi} + I_3 \dot{\psi} \cos \theta \equiv I_1 b \end{aligned} \right\} \begin{array}{l} \text{These} \\ \text{define} \\ a \text{ and } b \end{array}$$

Conservative system $\Rightarrow$ energy is conserved:

$$E = T + V = \frac{1}{2} I_1 (\dot{\theta}^2 + \dot{\phi}^2 \sin^2 \theta) + \underbrace{\frac{1}{2} I_3 \omega_3^2}_{\text{const.}} + Mgl \cos \theta$$

From P_ψ : $I_3 \dot{\psi} = I_1 a - I_3 \dot{\phi} \cos \theta$

Insert this in expression for P_ϕ :

$$I_1 \dot{\phi} \sin^2 \theta + I_1 a \cos \theta = I_1 b \quad [I_3 \text{ terms cancel}]$$

$$\Rightarrow \dot{\phi} = \frac{b - a \cos \theta}{\sin^2 \theta}; \quad \dot{\psi} = \frac{I_1 a}{I_3} - \frac{b - a \cos \theta}{\sin^2 \theta} \cdot \cos \theta$$

$$\Rightarrow E' \equiv E - \frac{I_3}{2} \omega_3^2 = \frac{I_1}{2} \dot{\theta}^2 + \frac{I_1}{2} \frac{(b - a \cos \theta)^2}{\sin^2 \theta} + Mgl \cos \theta$$

A one-dimensional problem in the variable θ with an effective potential

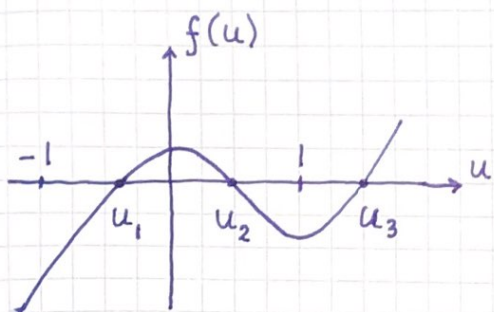
$$V'(\theta) = Mgl \cos \theta + \frac{1}{2} I_1 \left(\frac{b - a \cos \theta}{\sin \theta} \right)^2$$

The substitutions $u = \cos \theta$, $\alpha = 2E'/I_1$ and $\beta = 2Mgl/I_1$ and some algebra yields the equation of motion

$$\dot{u}^2 = \beta u^3 - (\alpha + a^2) u^2 + (2ab - \beta) u + (\alpha - b^2) \equiv f(u) \geq 0$$

Solution of $\dot{u}^2 = f(u)$ yields $\theta(t)$, the nutation of the x_3 axis relative to the vertical space axis x_3' . Next, $\theta(t)$ determines $\dot{\phi}(t)$, the precession of x_3 around x_3' , and $\dot{\psi}(t)$, the spinning around x_3 .

(72)



Solutions of $f(u)=0$
corresponds to $\dot{u}=0$
or $\dot{\Theta} = -\dot{u}/\sin\Theta = 0$, i.e.,
values of Θ where $\dot{\Theta}$
changes sign.

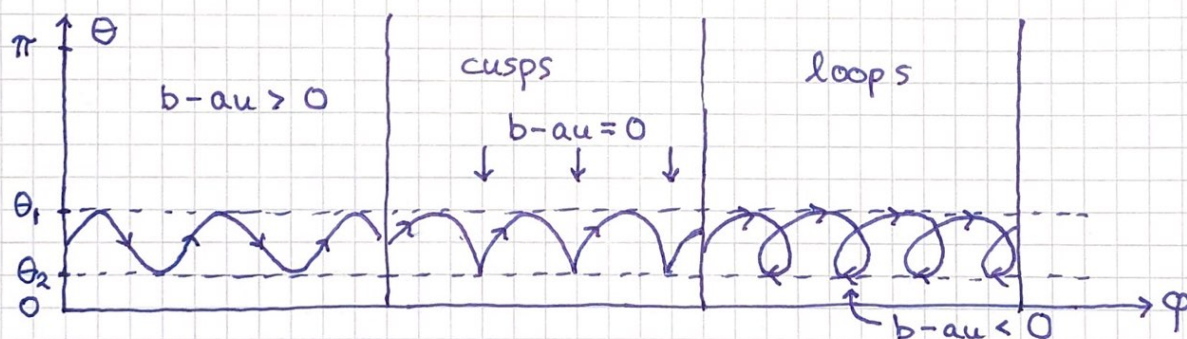
$$\left. \begin{aligned} f(\pm 1) &= -(b \mp a)^2 < 0 \\ f(\pm \infty) &= \pm \infty \end{aligned} \right\} \text{always a solution } u_3 > 1 \text{ which is not physical}$$

Acceptable values of $u = \cos\Theta$: $u_1 < u < u_2$

$$\text{Precession: } \dot{\varphi} = \frac{b - a \cos\Theta}{\sin^2\Theta} = \frac{b - au}{1 - u^2}$$

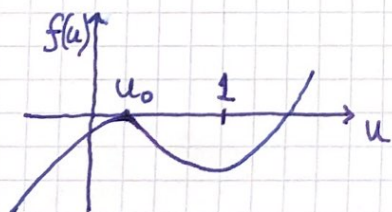
Usually $\dot{\Psi} \gg \dot{\varphi}$ for a spinning top.

May have nutation, $\Theta_2 < \Theta < \Theta_1$, with $\dot{\varphi} > 0$ always,
 $\dot{\varphi} \geq 0$ (cusps) or $\dot{\varphi} \leq 0$ (loops):



Regular precession: $\Theta(t) = \Theta_0 = \text{const.}$

$$\Rightarrow f(u) = 0 \text{ for } u = u_1 = u_2 = u_0 = \cos\Theta_0$$



$f=0$ and $df/du=0$
in u_0 (starting point for
a deeper analysis of regular prec.)